Performance of DoFP Polarimeter Calibration

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Abstract

Division-of-focal plane (DoFP) imaging polarimeters are useful instruments for measuring polarization images for a variety of applications, however only recent advances in nanofabrication have enabled the practical manufacture of these sensors. These sensors are made by integrating nanowire polarization filters directly with an imaging array, and size variations of the nanowires due to fabrication can cause the optical properties of the filters to vary up to 20% across the imaging array. If left unchecked, these variations introduce significant errors when reconstructing the polarization image. Calibration methods offer a means to correct these errors. This work evaluates a scalar and matrix calibration derived from a mathematical model of the polarimeter behavior. The methods are evaluated quantitatively with an existing DoFP polarimeter while varying illumination intensity and angle of linear polarization.

Keywords: polarization, polarimeter, division-of-focal-plane, scalar calibration, matrix calibration

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1. Introduction

Polarization, along with intensity and wavelength, is one of the fundamental properties of light. In essence, it describes the orientation of the light wave as it propagates through space. While humans are not capable of directly seeing the polarization of light, there are a variety of applications for recording images that include polarization information. These include remote sensing applications

such as the removal of atmospheric haze from photographs [Shwartz06], or enhancing contrast beyond that of intensity and spectral images [Tyo06]. Additionally, there are many animals that both display polarization patterning on their bodies and can directly see the polarization of light. Prominent examples include scarab beetles [Goldstein06], mantis shrimp and many cephalopods [Cronin03]. In order for biologists to effectively study these animals they must use imaging polarimeters.

In general, imaging polarimeters measure the polarization state of light across a plane. This is done by modulating the polarized light into components that can be captured with conventional intensitymeasuring image sensors and then reconstructing the original polarization image from its components. Popular modulation schemes include the division-of-time (DoT) polarimeter, where a filter wheel switches between polarization filters placed in front of the imager; the division-of-aperture (DoA) polarimeter where a prism splits the image into multiple paths, each of which has its own polarization optics and imager; and the division-of-focal-plane (DoFP) polarimeter, which splits the image by placing a repeating pattern of micro-filters directly onto the pixels of an imager. [Tyo06]

DoFP Polarimeters have several advantages over competing imaging polarimeter architectures. Most notably, they capture all of the components needed to reconstruct the incident polarization state simultaneously—which avoids the motion-blur inherent to DoT polarimeters. In addition, their monolithic design makes them more compact and robust than either DoT or DoA polarimeters, which makes them the ideal choice for field work.

However, they also have some notable sources of error. Instantaneous field of view (IFOV) errors have received the most attention from the research community. These result from the spatial modulation the filter array introduces to the image. Image interpolation and Fourier transform techniques have been developed to minimize IFOV errors [Tyo09][Gao11]. The second major source of error is non-uniformity and non-ideality in both the imaging and filter arrays. In both arrays, these non-idealities are caused by flaws introduced during the manufacturing process. Correcting these errors so that the imager behaves as close to the ideal as possible is known as "calibration," and the focus of this work will be the analysis and comparison of two calibration techniques. First we will examine the mathematic model of DoFP polarimeters and how the two calibration techniques are derived from it. Then we will examine the results of applying the techniques.

2. DoFP Polarimeter Model

As stated previously, division-of-focal-plane polarimeters are constructed by integrating an array of polarization filters with an array of imaging elements to create an array of polarization-sensitive pixels. Figure 1 shows the typical configuration of an array of linear polarization filters oriented at 0°, 45°, 90°, and 135°; and grouped into 2-by-2 squares called super-pixels. The model derived below, however, will apply to arbitrary polarization filters and super-pixel groupings. Each filter is aligned above a single intensity-measuring pixel.

In general, we can model the behavior of each polarization pixel as the composition of the pixel's conversion function acting on the intensity of the light passed by the polarization filter. Using Mueller calculus, the light passed by the polarization filter is represented by the Stokes vector \vec{S}_f as a function of the Stokes vector of the incident light, \vec{S}_{in} :

$$
\vec{S}_f = M_f \vec{S}_{in}
$$

where M_f is the 4×4 real-valued Mueller matrix that characterizes the filter [Goldstein11]. The intensity component of the filtered Stokes vector is then passed to the conversion function of the underlying pixel, which we will assume is linear and noiseless:

$$
I_p = g(1 \quad 0 \quad 0 \quad 0) \cdot \vec{S}_f + d = g(1 \quad 0 \quad 0 \quad 0)M_f \vec{S}_{in} + d,
$$

where I_p is the digital value measured by the pixel; g is the conversion gain of the pixel, which converts the intensity to digital values; and \boldsymbol{d} is the dark offset of the pixel. The dot product $(1 \ 0 \ 0 \ 0) \cdot \vec{S}_f$ selects only the intensity component of the Stokes vector. This model can be simplified somewhat by noting that only the first row of the Mueller matrix is required and we can combine the conversion gain into the dot product. Thus, we let

$$
\vec{A} = g(1 \quad 0 \quad 0 \quad 0)M_f,
$$

and the polarization pixel model becomes

$$
I_p = \vec{A} \cdot \vec{S}_{in} + d.
$$

The vector \vec{A} is known as the pixel's analysis vector. Note that the model does not include temporal noise, quantization noise, or saturation effects—they are beyond the scope of this work.

When considering a super-pixel of polarization pixels, we group the individual pixel responses into a vector so that

$$
\vec{I}_{sp} = \begin{pmatrix} \vec{A}_1 \cdot \vec{S}_{in} + d_1 \\ \vdots \\ \vec{A}_n \cdot \vec{S}_{in} + d_n \end{pmatrix} = \underbrace{\begin{pmatrix} \vec{A}_1 \\ \vdots \\ \vec{A}_n \end{pmatrix}}_{\mathbf{A}} \cdot \vec{S}_{in} + \underbrace{\begin{pmatrix} d_1 \\ \vdots \\ d_n \end{pmatrix}}_{\mathbf{d}} = \mathbf{A} \vec{S}_{in} + \vec{d},
$$

where **A** is the super-pixel's analysis matrix, which is constructed by stacking the analysis vectors \vec{A}_k for each pixel \bf{k} in the super-pixel. Note that this model assumes that the illumination is either uniform over the super-pixel or that all of the pixels in a super-pixel are co-located; this is what causes the IFOV reconstruction errors that are dealt with in [Tyo09].

3. Calibration Techniques

In the ideal case, all of the polarization pixels would have identical and perfect analysis vectors and dark offsets, \vec{A} and \vec{a} , respectively. Additionally, this would lead to all of the super-pixels having identical and perfect analysis matrices, \bm{A} . As mentioned previously, this is not typically the case and deviations from the ideal can introduce significant reconstruction errors. In light of this, the purpose of DoFP polarimeter calibration techniques is to transform the non-ideal polarization-pixel responses into the ideal polarization-pixel response.

In general, calibration techniques consist of a calibration function that takes a single pixel's intensity measurement or super-pixel's intensity measurement vector and returns an approximation of the ideal intensity value or vector that the pixel should have measured:

$$
I_{p,cal} = Cal_p(I_p) \otimes I_{p,ideal} = \vec{A}_{ideal} \cdot \vec{S}_{in}
$$
 for single pixels;

$$
\vec{I}_{sp,cal} = Cal_{sp}(\vec{I}_{sp}) \otimes \vec{I}_{sp,ideal} = A_{ideal} \vec{S}_{in}
$$
, for super-pixels.

The ideal intensity value or vector comes from the same model, but with some ideal value for the analysis vectors and zero dark offsets.

3.1 Scalar Calibration

The scalar, or "gain and offset," calibration method presented in [York11] arises when each pixel is considered independently. Since we are trying to transform the linear polarization pixel response into the similarly linear ideal pixel response, we use a linear transformation as the calibration function:

$$
I_{p,cal} = Cal_p(I_p) = g_c(I_p - d_c).
$$

The parameters g_c and d_c are the calibration gain and calibration dark value, respectively, and are determined such that they minimize the squared L2, or Euclidian, distance between $I_{p,cal}$ and $I_{p,ideal}$:

$$
\begin{array}{l}\n\min_{g_c, d_c} \left\| I_{p, cal} - I_{p, ideal} \right\|_2^2 \\
\min_{g_c, d_c} \left\| g_c \left(\vec{A} \cdot \vec{S}_{in} + d - d_c \right) - \vec{A}_{ideal} \cdot \vec{S}_{in} \right\|_2^2\n\end{array}
$$

Completing the minimization by setting the partial derivatives with respect to g_c and d_c equal to zero and solving the system of equations yields

$$
d_c = some\ constant,\ and\ g_c = \frac{\vec{A}_{ideal} \cdot \vec{S}_{in}}{d - d_c + \vec{A} \cdot \vec{S}_{in}}.
$$

By letting $d_c = d$ this simplifies to

$$
d_c = d, \text{ and } g_c = \frac{\vec{A}_{ideal} \cdot \vec{S}_{in}}{\vec{A} \cdot \vec{S}_{in}}.
$$

Unfortunately, the presence of \vec{S}_{in} in the expression for g_c means that the optimal calibration gain depends on the light we are trying to measure. This is clearly unacceptable since we cannot accurately measure the light without an appropriate value for both g_c and d_c .

One solution to this problem is to make the approximation that \vec{A} is a scalar multiple of \vec{A}_{ideal} . Then $(\vec{A}_{ideal} \cdot \vec{S}_{in})/(\vec{A} \cdot \vec{S}_{in})$ is approximately constant and we can simplify g_c to the ratio of the length of \vec{A} to that of \vec{A}_{ideal} :

$$
d_c = d, \text{ and } g_c = \frac{\left\| \vec{A}_{ideal} \right\|_2}{\left\| \vec{A} \right\|_2}
$$

This is equivalent to the expressions given in equations 7, 8, and 9 of [York11] if we take the ideal response as that of the best-performing pixel of the super-pixel rotated to match the orientation of the current pixel.

3.2 Matrix Calibration

When we consider each super-pixel as a whole, we can derive the calibration function presented in [Myhre12]. This calibration function is similar to the scalar version, but the calibration dark value is expanded to a vector, and the calibration gains are a square matrix:

$$
\vec{l}_{sp,cal} = Cal_{sp}(\vec{l}_{sp}) = G_c(\vec{l}_{sp} - \vec{d}_c).
$$

Again, G_c and \vec{d}_c are determined such that they minimize the squared L2 distance between $\vec{l}_{sp,cal}$ and $\vec{l}_{sp.idgal}$:

$$
\begin{array}{l}\n\min_{G_O, d_C} \left\| \vec{I}_{sp, cal} - \vec{I}_{sp, ideal} \right\|_2^2 \\
\min_{G_O, d_C} \left\| G_c \left(A \vec{S}_{in} + \vec{d} - \vec{d}_c \right) - A_i \vec{S}_{in} \right\|_2^2\n\end{array}
$$

Completing this minimization by setting the partial derivatives equal to zero and solving the resulting system of equations leads to

$$
\vec{d}_c = \vec{d} \text{ and } \mathbf{G}_c = \mathbf{A}_i \mathbf{A}^{-1},
$$

in the case where \bf{A} is square and full rank or

 $\vec{d}_c = \vec{d}$ and $\vec{G}_c = A_i A^T (AA^T)^{-1}$

otherwise. While these results are very similar to the scalar case, this calibration function is clearly more general and powerful as it does not include \vec{S}_{in} in the expressions for \vec{d}_c or \vec{G}_c and it makes no further assumptions about the polarization pixels.

3.3 Geometric Interpretation

The polarization pixel model $I_p = \vec{A} \cdot \vec{S}_{in} + d$ can be interpreted as projecting the \vec{S}_{in} vector onto the \vec{A} vector and applying some offset. Applying the scalar calibration function results in a new projection

$$
I_{p,cal} = g_c(\vec{A} \cdot \vec{S}_{in} + d - d_c) = \frac{\|\vec{A}_{ideal}\|_2}{\|\vec{A}\|_2} \vec{A} \cdot \vec{S}_{in} \approx \vec{A}_{ideal} \cdot \vec{S}_{in}
$$

where we have removed the offset and rescaled the analysis vector to the length of the ideal analysis vector. Thus the accuracy of this method depends on the actual and ideal analysis vectors pointing in the same direction.

Alternatively, the matrix calibration function multiplies each of the analysis vectors by a matrix:

$$
\vec{l}_{sp,cal} = G_c(A\vec{S}_{in} + \vec{d} - \vec{d}_c) = G_cA\vec{S}_{in} \otimes A_{ideal}\vec{S}_{in}.
$$

As long as the analysis matrix has high enough rank, the gain matrix will be able to both scale and rotate the individual analysis vectors to match the ideal case. Because of this ability to rotate the analysis vectors, the matrix calibration technique should achieve significantly lower error rates than the scalar technique.

4. Evaluation

4.1 Experimental Setup

The two calibration functions were tested on a 300x300 pixel sub-region of the linear DoFP polarimeter described in [Gruev10]. This polarimeter is composed of linear pixels aligned with a repeating pattern of 2x2 blocks of linear polarization filters oriented at 0°, 45°, 90°, and 135°. Because it only has linear polarization filters and no circular polarization filters, the last component of its pixels' analysis vectors is always zero, and it can only reconstruct the first three parameters of the Stokes vector.

Performing the evaluation requires two steps. The first is to determine the analysis vectors for each pixel. Since the pixel model is linear, we can simply measure the pixel responses under a variety of illuminations and use a multivariate linear regression to solve for \vec{A} and \vec{a} for each pixel. The second step is to then calculate the appropriate calibration parameters for each function and evaluate their performance. The apparatus used to illuminate the polarimeter is shown in figure 2. By varying the adjustable shutter, the intensity of the illumination, S_0 , can be controlled. Varying the angle of the

linear polarizer, θ , allows us to generate values for S_1 and S_2 on a circle with a radius of S_0 . The calibrated photodiode lets us measure the absolute level of S_0 . This apparatus leaves the last value of the Stokes vector, S_3 , at zero, but since the polarimeter we are evaluating cannot measure S_3 it is inconsequential.

Figure 2: Apparatus for evaluating the calibration functions. It generates uniform illumination with arbitrary values of S_0 , and S_1 , S_2 on a circle of radius S_0 .

4.2 Pixel Model Regression

The 300x300 pixel window was illuminated with 6 different intensities of light, ranging from 2.5% to 100% of the dynamic range of the polarimeter; and 9 different angles, equally spaced around the S_1 , S_2 circle. One hundred samples were taken at each intensity-angle pair to reduce the effects of temporal noise on the regression. This results in 5,400 samples taken for each of the 90,000 pixels. Each pixel's analysis vector and dark offset were solved for using a basic multilinear regression. The histogram in figure 3 (a) summarizes the coefficents of determination, \mathbb{R}^2 values, for all of the regressions, which shows that the regressions are not perfect, but do explain at least 99.7% of the variation in the sample data. Plot (b) shows the reconstruction errors versus the predicted values, which has a slight trend, but not significant compared to the range of the data. Plot (c) shows a normal quantile-quantile plot of the errors—this indicates that the error distribution has longer tails than the normal distribution, but is otherwise closely approximated by it. Finally, plot (d) shows a histogram of the ratio of each regression's F-value to the 99% significance F-value, all of which are much greater than 1, which indicates that all of the regressions are significant.

Figure 3: Statistics of the pixel model regression: (a) coefficients of determination, \mathbb{R}^2 ; (b) reconstruction error vs. reconstructed pixel response; (c) normal quantile-quantile plot; and (d) F-tests.

4.3 Calibration Function Performance

Both scalar and matrix gains were computed from the analysis vectors solved for with the linear regression. Figure 4 shows plots of the pixels analysis vectors and the effects of applying the calibration functions to the analysis vectors. Plots (a) and (b) show the actual analysis vectors of the pixels. They have a clear error relative to the ideal but little spread in the normalized A1-A2 space shown in (a), but very large spread in the A0 dimension. Applying the scalar gains results in plots (c) and (d). As expected, there is no change in the normalized A1-A2 space, but the variation in A0 is reduced dramatically. Finally, multiplying the vectors by their respective matrix gains results in plots (e) and (f). Since this calibration function is powerful enough to meet the ideal without any extra assumptions, the effective analysis vectors are transformed completely to their ideal values.

Figure 4: Plots of the effective analysis vectors after calibration. Plots (a) and (b) show the pixels' actual analysis vectors, (c) and (d) show the effect of applying scalar-calibration, and (e) and (f) show the effect of applying matrix-calibration. Plots (a), (c), and (e) show the vectors in the A1-A2 space, normalized by A0. The radius in these plots indicates how well the filters discriminate different polarizations of light. Plots (b), (d), and (f) have the same x-axis, but show A0 on the y-axis. These plots indicate the overall transmission ratio of the filters. In all plots the ideal values are indicated by red squares.

Figure 5 examines the results of applying the calibration functions to the actual pixel responses. Histogram (a) compares all of the super-pixels' 0° components uncalibrated response and the results of scalar and matrix calibrations with the ideal response. The uncalibrated response has a very large spread, which shows the non-uniformity of the filters' analysis vectors. Applying scalar calibration reduces this spread significantly, which indicates that the non-uniformity is predominantly in the A_0 coefficients of the filters, but does not center the responses about the ideal value. Finally, the matrix calibration further reduces the spread and centers the responses about the ideal, which indicates that the matrix calibration function accounts for significantly more of the non-uniformities and nonidealities in the pixel responses. This agrees with the conclusions drawn from figure 4. Histograms (b), (c), and (d) respectively show the uncalibrated, scalar-calibrated, and matrix-calibrated responses of all of the polarization pixels when illuminated with uniform linearly polarized light at 15°. In histogram (b), it is clear that the spread of the responses is heavily dependent on the response level. This is further evidence that the majority of the non-uniformities of the filters is in the A_0 coefficients. Further support for this is seen in the lack of this dependence in the scalar-calibrated responses, which predominantly corrects A_0 .

Figure 5: Histograms of linear polarization pixel responses when un-calibrated (a) and (b), scalarcalibrated (a) and (c), and matrix-calibrated (a) and (d) compared to the ideal response.

Finally, figure 6 shows the mean and standard deviation of the un-calibrated, scalar-calibrated, and matrix-calibrated responses to linearly polarized light as the S_1 and S_2 components of the incident light are swept about the circle with radius S_0 . The plots of figure 6 support all of the conclusions

drawn from the histograms of figure 5 as the incident angle of polarization changes. We can see that as the incident angle changes, the un-calibrated mean responses of the different orientations of polarization pixels have a sinusoidal response, but have non-uniform amplitudes and offsets. Both the scalar- and the matrix-calibrated data correct a majority of this non-uniformity, however it is worth noting that there is a significant offset from zero in the scalar-calibrated responses, and that the peaks of the scalar-calibrated curves are not centered on the appropriate angles. The matrix-calibrated responses clearly corrects for both problems.

Figure 6: Mean and standard deviations of un-calibrated (a), scalar-calibrated (b), and matrixcalibrated (c) polarization pixel responses with varying incident angle of polarization.

5. Summary

We have shown two DoFP polarimeter calibration functions, one based on treating each pixel individually, and the other treating each super-pixel as a whole. Both were derived from a linear, noiseless model for a polarization pixel and were evaluated on an existing DoFP polarimeter. The parameters of the linear model were learned for each pixel of the polarimeter with very high confidence by performing a linear regression over 5,400 test images. The results show that while the scalar calibration technique reduces the reconstruction error greatly, it is not as effective as the matrix-based technique. This was predicted from the derivation of the two methods as the scalar calibration function required making extra assumptions about the nature of the pixels, which were unnecessary in the matrix case.

List of Acronyms

DoFP: Division of Focal Plane **DoT:** Division of Time **DoA:** Division of Amplitude **IFoV:** Instantaneous Field of View

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